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REVISED AND EXTENDED ANALYSIS IN FOUR TIMES IONIZED XENON XE V

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Abstract—The atomic-emission spectrum of four times ionized xenon (Xe V) has been observed in the region 500–6700 Å; 233 lines were classified as transition between $5s5p^3$, $5s^25p5d$, $5s^25p6s$ with $5s^25p^2$, $5s^25p6p$ and $5s^25p4f$ configurations. Ninety-seven of these lines are reported for the first time. Two new energy level values corresponding to the $5s^25p4f$ configuration have been determined, and we proposed two new values for the previously reported 3D_1 and 3G_5 energy levels of this configuration. The value of the $5s^25p6p$ 1S_0 energy level has also been determined, and we propose a new value for the previously reported $5s^25p5d$ 3F_4 energy level. New adjusted values for the previously known levels of the studied configurations are included. Least-squares fitted parametric calculations involving configuration interactions have been carried out to interpret the observed spectrum. © 1998 Elsevier Science Ltd. All rights reserved.

1. INTRODUCTION

There is great interest in spectroscopy data from xenon due to applications in collision physics, high-temperature plasmas and laser physics. In this last point, the understanding of the population mechanisms that affect the laser action for this element, require a knowledge of the spectral analysis corresponding to the different ions involved.

The spectrum of four times ionized xenon (Xe V) has the ground configuration $5s^25p^2$, the same as Sn I. Spectra of Sb II,¹ I IV,² Cs VI³ and Ba VII⁴ in this sequence were previously analysed by other authors. Several ionic UV-visible Xe lines were charge-state assigned to Xe V by Duchowicz et al⁵ but not classifications. Gallardo et al⁶ obtained nine new levels of the $5s^25p6p$ configuration, the first established energy levels of an even excited configuration. Pinnington et al⁷ have presented a study of the beam-foil spectrum of Xe V. Recently, Tauheed et al⁸ made the analysis of several configurations of this ion using a gas-triggered spark as a light source. Very recently, a work concerning $5s^25p5d$ – $5s^25p4f$ and $5s5p^3$ – $5s^25p4f$ transitions, establishing 10 new levels of the $5s^25p4f$ configuration, the remaining 3F_4 level of $5s^25p5d$ and the classification of several known Xe V laser lines, was reported by Larsson et al,⁹ using collision-based spectroscopy.

With the final goal of completing the classification of the most prominent UV-visible Xe laser lines, we have continued an investigation of the Xe V spectrum that concerns the study of the energy levels in the configuration $5s^25p4f$, and the transitions $5s5p^3$ – $5s^25p4f$, $5s^25p5d$ – $5s^25p4f$, $5s^25p6s$ – $5s^25p4f$, and the classification of lines belonging to transitions between the $5s^25p6p$ configuration with odd parity configurations.

2. EXPERIMENT

The experimental set-up is similar to that described in the work of Gallardo et al⁶ SWR plates were used to record the new spectra in the vacuum ultraviolet region, where known lines of C, N, O, Xe II and Xe III were used as internal standards. In the UV-visible region new spectra were recorded using a water-cooled Pyrex tube 1.3 m long and 3 mm internal bore. The indium-coated tungsten

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electrodes were placed in side arms 1 m apart. Using a 140 nF capacitor, peak current pulses of 3.5 kA and 1 μ s (FWHM) were obtained. In certain cases, larger capacitances up to 280 nF were used. Thorium lines from an electrodeless discharge were superimposed on the spectrograms and served as reference lines.¹⁰

The spectrograms were measured with a photoelectric automatic Grant comparator, and the accuracy of the wavelength values is estimated to be $\pm 0.01 \text{ \AA}$ in the first diffraction order. The intensities of the lines were based on visual estimates. To distinguish among different states of ionization, we studied the behaviour of the spectral lines intensities as a function of pressure.

3. RESULT AND DISCUSSION

The lines observed in the present work, are listed in Table 1. The total number of observed lines is 233, of which 97 correspond to new classified lines. Table 1 also contains almost all the lines previously reported in the work of Duchowicz et al⁵ as laser lines. It is interesting to note that most of these laser lines correspond, in spontaneous emission, to high intensity transitions between the

Table 1. Classified lines in the Xe V spectrum

Intensity	$\lambda [\text{\AA}]$ (vac.)	$\sigma [\text{cm}^{-1}]$		Ref.	Classification	
		Obs	Calc			
10	543.05	184145.1	7.6	6	$5s^25p^2$	$^3P_0-5s^25p5d$ 3P_1
11	566.57	176500.7	3.2	6	$5s^25p^2$	$^3P_1-5s^25p5d$ 1D_2
8	571.90	174855.7	5.8	7,8	$5s^25p^2$	$^3P_1-5s^25p5d$ 3P_1
7	575.59	173734.7	3.4	8	$5s^25p^2$	$^3P_1-5s^25p5d$ 3P_0
6	577.79	173073.3	1.7	8	$5s^25p^2$	$^3P_0-5s^25p5d$ 3D_1
7	582.52	171667.9	8.3	8	$5s^25p^2$	$^3P_2-5s^25p5d$ 1D_2
6	582.93	171547.2	7.8	8	$5s^25p^2$	$^1D_2-5s^25p5d$ 1P_1
5	588.16	170021.8	0.9	8	$5s^25p^2$	$^3P_2-5s^25p5d$ 3P_1
4	593.22	168571.5	69.8	7,8	$5s^25p^2$	$^1S_0-5s^25p6s$ 1P_1
8	595.09	168040.9	0.5	7,8	$5s^25p^2$	$^3P_2-5s^25p5d$ 3D_3
1	599.24	166878.0	7.6	8	$5s^25p^2$	$^3P_2-5s^25p5d$ 3D_2
7	603.40	165727.5	6.8	7,8	$5s^25p^2$	$^1D_2-5s^25p5d$ 1F_3
6	610.57	163781.4	79.9	8	$5s^25p^2$	$^3P_1-5s^25p5d$ 3D_1
9	635.39	157383.7	3.8	7,8	$5s^25p^2$	$^1D_2-5s^25p5d$ 1D_2
8	637.51	156860.2	0.9	7,8	$5s^25p^2$	$^3P_2-5s^25p5d$ 3D_2
6	662.10	151034.5	3.7	6	$5s^25p^3$	$^5S_2-5s^25p6p$ 1P_1
7	679.28	147214.7	5.0	7,8	$5s^25p^2$	$^3P_1-5s^25p5d$ 3F_2
7	691.27	144661.3	0.5	8	$5s^25p^2$	$^1D_2-5s^25p5d$ 3D_1
6	701.38	142576.1	6.4	8	$5s^25p^2$	$^1D_2-5s^25p5d$ 3P_2
12	702.34	142381.2	0.1	8	$5s^25p^2$	$^3P_2-5s^25p5d$ 3F_2
12A	702.87	142273.8	2.8	6	$5s5p^3$	$^5S_2-5s^25p6p$ 3P_1
10	707.26	141390.7	1.4	7,8	$5s^25p^2$	$^3P_2-5s5p^3$ 3S_1
6	707.90	141263.0	1.4	8	$5s^25p^2$	$^1D_2-5s5p^3$ 1P_1
3	735.94	135880.6	2.1	6	$5s5p^3$	$^5S_2-5s^25p6p$ 3D_1
2	765.47	130638.7	8.2		$5s5p^3$	$^3D_3-5s^25p6p$ 1D_2
3	768.58	130110.1	1.0		$5s5p^3$	$^3D_2-5s^25p6p$ 3D_3
7	772.00	129533.7	5.0	6	$5s5p^3$	$^3D_1-5s^25p6p$ 3P_2
6 bl	776.87	128721.7	4.3	6	$5s5p^3$	$^3D_2-5s^25p6p$ 3P_2
6	777.61	128599.2	601.3	8	$5s^25p^2$	$^1S_0-5s^25p5d$ 3D_1
7	780.67	128095.1	5.6	7,8	$5s^25p^2$	$^1D_2-5s^25p5d$ 3F_2
3	786.66	127119.7	9.5		$5s5p^3$	$^3D_2-5s^25p6p$ 1P_1
8	791.84	126288.1	9.0	6	$5s5p^3$	$^3D_3-5s^25p6p$ 3D_3
5	797.38	125410.7	0.9	8	$5s^25p^2$	$^3P_1-5s5p^3$ 3P_2
7	798.19	125283.5	3.4	8	$5s^25p^2$	$^3P_1-5s5p^3$ 3P_1
5	798.71	125201.9	2.2	8	$5s^25p^2$	$^1S_0-5s5p^3$ 1P_1

continued

Table 1—continued

Intensity	$\lambda [\text{\AA}]$ (vac.)	$\sigma [\text{cm}^{-1}]$		Ref.	Classification	
		Obs	Calc			
7	805.70	124115.7	6.3	8	$5s^25p^2$	$^3P_1-5s5p^3$
9	829.35	120576.4	6.0	7,8	$5s^25p^2$	$^3P_2-5s5p^3$
6	830.23	120448.6	8.5	8	$5s^25p^2$	$^3P_2-5s5p^3$
1	834.08	119892.6	2.6	6	$5s5p^3$	$^3D_1-5s^25p6p$
2	839.15	119168.2	9.3		$5s5p^3$	$^3D_1-5s^25p6p$
6	844.89	118358.6	8.6	6	$5s5p^3$	$^3D_2-5s^25p6p$
7	851.82	117395.7	5.8	7,8	$5s^25p^2$	$^1D_2-5s5p^3$
6	862.20	115982.4	2.0	6	$5s5p^3$	$^3P_1-5s^25p6p$
3	863.15	115854.7	4.5		$5s5p^3$	$^3P_2-5s^25p6p$
8	867.40	115287.1	6.3	7,8	$5s^25p^2$	$^3P_0-5s5p^3$
9A	886.68	112780.3	78.6	6	$5s5p^3$	$^5S_2-5s^25p6p$
1	893.12	111967.0	7.9	6	$5s5p^3$	$^3D_2-5s^25p6p$
10	896.82	111505.1	5.3	6	$5s5p^3$	$^3P_2-5s^25p6p$
6	900.51	111048.2	7.7	8	$5s^25p^2$	$^1S_0-5s5p^3$
9	907.06	110246.3	6.1	6	$5s5p^3$	$^3P_1-5s^25p6p$
6	908.11	110118.8	8.6	6	$5s5p^3$	$^3P_2-5s^25p6p$
7	921.54	108514.0	3.8	6	$5s5p^3$	$^3P_2-5s^25p6p$
13	936.29	106804.5	5.2	7,8	$5s^25p^2$	$^3P_1-5s5p^3$
8	940.81	106291.4	1.5	7,8	$5s^25p^2$	$^1D_2-5s5p^3$
5	941.95	106162.7	4.0	6	$5s^25p^2$	$^1D_2-5s5p^3$
7	943.44	105995.1	4.5	8	$5s^25p^2$	$^3P_1-5s5p^3$
7	954.66	104749.3	50.2	6	$5s5p^3$	$^1D_2-5s^25p6p$
12	980.68	101970.1	0.3	6	$5s^25p^2$	$^3P_2-5s5p^3$
10	988.55	101158.3	9.6	7,8	$5s^25p^2$	$^3P_2-5s5p^3$
3	989.63	101047.9	7.5		$5s5p^3$	$^3P_0-5s^25p6p$
5	995.26	100476.3	6.2		$5s5p^3$	$^3P_2-5s^25p6p$
10	1002.48	99752.6	2.9		$5s5p^3$	$^3P_2-5s^25p6p$
9	1026.60	97408.9	9.5	6	$5s5p^3$	$^1D_2-5s^25p6p$
3	1056.47	94654.8	6.8	6	$5s5p^3$	$^3P_0-5s^25p6p$
6	1059.34	94398.4	8.7		$5s5p^3$	$^3D_3-5s^25p4f$
4	1063.26	94050.4	0.4	6	$5s^25p5d$	$^3F_2-5s^25p6p$
7	1063.55	94024.7	4.4		$5s5p^3$	$^3D_1-5s^25p4f$
8	1071.10	93362.0	2.2		$5s5p^3$	$^3P_2-5s^25p6p$
6	1072.81	93213.2	3.7		$5s5p^3$	$^3D_2-5s^25p4f$
3	1083.53	92290.9	2.3	6	$5s5p^3$	$^3S_1-5s^25p6p$
4	1086.01	92080.1	0.4		$5s5p^3$	$^3D_1-5s^25p4f$
6	1092.79	91508.9	7.8	7,8	$5s5p^2$	$^1D_2-5s5p^3$
8	1095.65	91270.0	69.7		$5s5p^3$	$^3D_2-5s^25p4f$
6	1109.81	90105.5	4.8	8	$5s^25p^2$	$^1S_0-5s5p^3$
5	1111.49	89969.3	9.7		$5s5p^3$	$^1P_1-5s^25p6p$
3	1114.82	89700.6	1.2	6	$5s^25p5d$	$^3F_2-5s^25p6p$
1	1115.28	89663.6	1.8		$5s5p^3$	$^3D_2-5s^25p4f$
6	1118.92	89371.9	1.9	6	$5s5p^3$	$^1D_2-5s^25p6p$
4	1119.77	89304.1	3.2	6	$5s5p^3$	$^3S_1-5s^25p6p$
3	1128.06	88647.8	8.6	6	$5s5p^3$	$^1D_2-5s^25p6p$
5	1132.32	88314.3	4.5	6	$5s^25p5d$	$^3F_2-5s^25p6p$
6	1140.28	87697.8	8.4	6	$5s5p^3$	$^3S_1-5s^25p6p$
8	1143.53	87448.5	7.7		$5s5p^3$	$^3D_3-5s^25p4f$
6	1153.27	86710.0	9.7	6	$5s^25p5d$	$^3F_2-5s^25p6p$
4	1155.13	86570.3	0.6		$5s^25p5d$	$^3F_2-5s^25p6p$
11	1164.96	85839.9	9.8	9	$5s5p^3$	$^3D_3-5s^25p4f$
9	1168.52	85578.3	7.6		$5s^25p5d$	$^3F_3-5s^25p6p$
10	1191.69	83914.4	3.2	9	$5s5p^3$	$^3D_2-5s^25p4f$
7	1218.88	82042.5	2.9		$5s5p^3$	$^3P_2-5s^25p4f$
12	1225.10	81626.0	6.2	9	$5s5p^3$	$^3D_3-5s^25p4f$
4	1236.34	80883.9	4.6	6	$5s5p^3$	$^1P_1-5s^25p6p$
7	1248.58	80091.0	1.2		$5s5p^3$	$^3D_3-5s^25p4f$
5	1255.32	79661.0	0.8	6	$5s5p^3$	$^3S_1-5s^25p6p$

continued

Table 1—continued

Intensity	$\lambda [\text{\AA}]$ (vac.)	$\sigma [\text{cm}^{-1}]$		Ref.	Classification	
		Obs	Calc			
6	1256.75	79570.3	69.6	6	$5s^25p5d$	$^3P_2-5s^25p6p$
6	1271.10	78672.0	2.1	6	$5s^25p5d$	$^3F_2-5s^25p6p$
12	1281.13	78056.1	6.1	7,8	$5s^25p^2$	$^3P_2-5s5p^3$
12	1282.89	77949.0	8.8	6	$5s^25p5d$	$^3F_2-5s^25p6p$
3	1290.57	77485.1	5.5	6	$5s^25p5d$	$^3D_1-5s^25p6p$
4	1301.69	76823.2	2.8	6	$5s^25p5d$	$^3P_2-5s^25p6p$
12	1317.46	75903.6	26		$5s5p^3$	$^3P_0-5s^25p4f$
9	1324.38	75507.0	6.5	9	$5s5p^3$	$^3D_2-5s^25p4f$
4	1324.58	75495.6	4.7		$5s^25p5d$	$^3P_1-5s^25p6p$
4	1329.44	75219.6	20.4	6	$5s^25p5d$	$^3P_2-5s^25p6p$
3	1330.70	75148.4	8.7	6	$5s5p^3$	$^1P_1-5s^25p6p$
10	1338.05	74735.6	5.5		$5s5p^3$	$^3P_1-5s^25p4f$
5	1340.34	74607.9	8.0		$5s5p^3$	$^3P_2-5s^25p4f$
12	1341.42	74547.9	8.5	6	$5s^25p5d$	$^3F_3-5s^25p6p$
7	1354.40	73833.4	3.7	6	$5s^25p5d$	$^3P_2-5s^25p6p$
13	1359.31	73566.7	6.8	9	$5s5p^3$	$^3D_2-5s^25p4f$
4	1359.74	73543.5	3.9	6	$5s5p^3$	$^1P_1-5s^25p6p$
11	1373.79	72791.3	1.5		$5s5p^3$	$^3P_1-5s^25p4f$
10	1376.20	72663.9	4.0		$5s5p^3$	$^3P_2-5s^25p4f$
6	1393.73	71749.9	9.6	6	$5s^25p5d$	$^3D_1-5s^25p6p$
9A	1397.45	71558.9	8.1	6	$5s^25p5d$	$^3F_2-5s^25p6p$
9	1407.34	71056.0	6.1	9	$5s5p^3$	$^3P_2-5s^25p4f$
7	1409.67	70938.6	8.6		$5s5p^3$	$^1D_2-5s^25p4f$
10	1412.12	70815.5	5.9	6	$5s^25p^2$	$^1S_0-5s5p^3$
12	1413.91	70725.9	5.6	9	$5s5p^3$	$^3D_3-5s^25p4f$
3	1425.61	70145.4	4.8	6	$5s^25p5d$	$^3D_1-5s^25p6p$
1	1433.80	69744.7	4.8		$5s5p^3$	$^3D_3-5s^25p4f$
1	1437.75	69553.1	2.9	6	$5s^25p5d$	$^3D_2-5s^25p6p$
8	1496.38	66827.9	7.7	9	$5s5p^3$	$^3D_3-5s^25p4f$
9	1505.80	66409.9	9.6	6	$5s^25p5d$	$^3P_1-5s^25p6p$
12	1531.22	65307.4	7.5	9	$5s5p^3$	$^3P_2-5s^25p4f$
4	1533.66	65203.5	3.7	6	$5s^25p5d$	$^3D_2-5s^25p6p$
6	1543.61	64783.2	3.0	6	$5s5p^3$	$^1P_1-5s^25p6p$
9	1544.11	64762.2	2.2	6	$5s^25p5d$	$^1D_2-5s^25p6p$
6	1557.84	64191.4	1.3	6	$5s^25p5d$	$^3P_2-5s^25p6p$
14	1561.47	64042.2	0.8		$5s^25p5d$	$^3D_3-5s^25p6p$
7	1566.96	63817.8	7.0	6	$5s^25p5d$	$^3D_2-5s^25p6p$
5	1568.10	63771.4	1.6	6,7	$5s^25p^2$	$^1D_2-5s5p^3$
7	1574.73	63502.9	3.7		$5s5p^3$	$^1D_2-5s^25p4f$
8	1575.59	64468.3	8.0	6	$5s^25p5d$	$^3P_2-5s^25p6p$
7	1607.41	62211.9	2.2	6	$5s^25p5d$	$^3D_2-5s^25p6p$
3	1610.11	62107.6	7.2	6	$5s^25p5d$	$^3D_1-5s^25p6p$
8	1624.41	61560.8	59.7		$5s5p^3$	$^1D_2-5s^25p4f$
2	1629.08	61384.3	3.9	6	$5s^25p5d$	$^3D_1-5s^25p6p$
6	1648.14	60674.5	3.7	6	$5s^25p5d$	$^3P_1-5s^25p6p$
8	1655.26	60413.5	3.0		$5s^25p5d$	$^1D_2-5s^25p6p$
12	1668.00	59952.0	1.8	9	$5s5p^3$	$^1D_2-5s^25p4f$
10	1675.51	59683.3	3.3		$5s^25p5d$	$^1P_1-5s^25p6p$
7	1692.91	59069.9	8.9		$5s^25p5d$	$^3P_1-5s^25p6p$
10	1712.52	58393.5	2.3		$5s5p^3$	$^1P_1-5s^25p6p$
3	1741.50	57421.8	1.5		$5s^25p5d$	$^1D_2-5s^25p6p$
6	1753.52	57028.1	8.3		$5s5p^3$	$^3P_1-5s^25p4f$
6	1775.44	56324.1	4.5		$5s^25p6s$	$^3P_1-5s^25p6p$
8	1782.05	56115.1	5.2		$5s^25p5d$	$^3F_3-5s^25p4f$
5	1818.40	54993.4	3.2		$5s^25p5d$	$^3D_1-5s^25p6p$
9	1844.91	54203.2	3.2	9	$5s5p^3$	$^1D_2-5s^25p4f$
5	1845.87	54175.0	4.6		$5s^25p5d$	$^3D_2-5s^25p6p$
6	1862.62	53687.8	7.3	9	$5s^25p5d$	$^3F_3-5s^25p4f$

continued

Table 1—continued

Intensity	$\lambda [\text{\AA}]$ (vac.)	$\sigma [\text{cm}^{-1}]$		Ref.	Classification	
		Obs	Calc			
4	1870.87	53451.1	1.3		5s ² 5p5d	³ D ₂ -5s ² 5p6p
3	1893.79	52804.2	3.9		5s ² 5p5d	³ F ₂ -5s ² 5p4f
2	1920.49	52070.0	0.0		5s ² 5p5d	¹ F ₃ -5s ² 5p6p
10	1921.44	52044.3	4.0	9	5s5p ³	³ P ₂ -5s ² 5p4f
3	1928.68	51848.9	8.6		5s5p ³	³ S ₁ -5s ² 5p4f
2	1944.34	51431.3	0.4		5s ² 5p5d	³ P ₀ -5s ² 5p6p
4	1959.60	51030.8	1.3		5s ² 5p5d	³ P ₁ -5s ² 5p6p
6	1976.36	50598.1	8.2		5s ² 5p5d	¹ P ₁ -5s ² 5p6p
5	1987.79	50307.1	8.0		5s ² 5p5d	³ P ₁ -5s ² 5p6p
$\lambda (\text{\AA})$ (air)						
2	2024.35	49382.7	3.9		5s ² 5p5d	¹ D ₂ -5s ² 5p6p
8	2029.75	49251.3	2.0		5s ² 5p5d	³ F ₂ -5s ² 5p4f
6	2032.56	49183.2	3.4		5s ² 5p6s	³ P ₀ -5s ² 5p6p
5	2040.86	48983.2	3.6		5s ² 5p6s	³ P ₁ -5s ² 5p6p
1	2089.12	47851.8	1.4		5s ² 5p5d	¹ P ₁ -5s ² 5p6p
8	2123.69	47073.0	3.0		5s5p ³	¹ P ₁ -5s ² 5p4f
4	2124.25	47060.6	0.6		5s ² 5p5d	³ D ₂ -5s ² 5p6p
5	2138.99	46736.3	6.3		5s ² 5p5d	³ F ₃ -5s ² 5p4f
9	2145.16	46601.9	2.1		5s ² 5p6s	¹ P ₁ -5s ² 5p6p
8	2184.69	45758.8	8.0		5s ² 5p5d	³ P ₂ -5s ² 5p4f
3	2228.33	44862.7	2.3	6	5s ² 5p5d	¹ P ₁ -5s ² 5p6p
4	2311.04	43257.3	7.5	6	5s ² 5p5d	¹ P ₁ -5s ² 5p6p
11	2409.59	41488.3	8.3	6	5s ² 5p6s	³ P ₂ -5s ² 5p6p
2	2435.82	41041.5	0.9		5s ² 5p5d	¹ F ₃ -5s ² 5p6p
14bl	2441.49	40946.2	6.0		5s ² 5p6s	³ P ₁ -5s ² 5p6p
15	2441.86	40940.0	39.7		5s5p ³	¹ D ₂ -5s ² 5p4f
10	2443.37	40914.7	4.8	9	5s ² 5p5d	³ F ₃ -5s ² 5p4f
12	2473.11	40422.7	2.5	6	5s ² 5p6s	³ P ₀ -5s ² 5p6p
9	2485.41	40222.7	2.7	6	5s ² 5p6s	³ P ₁ -5s ² 5p6p
5	2513.88	39767.1	6.4	6	5s ² 5p6s	³ P ₀ -5s ² 5p6p
1	2522.04	39638.6	8.1		5s5p ³	¹ P ₁ -5s ² 5p4f
11	2538.60	39380.0	79.8	9	5s ² 5p5d	³ F ₃ -5s ² 5p4f
6	2664.68	37516.8	7.0	6	5s ² 5p6s	¹ P ₁ -5s ² 5p6p
15w	2691.74	37139.7	9.1	5,6	5s ² 5p6s	³ P ₂ -5s ² 5p6p
3	2780.10	35959.4	9.4		5s ² 5p5d	³ F ₄ -5s ² 5p4f
6	2796.17	35752.7	2.4	6	5s ² 5p6s	³ P ₂ -5s ² 5p6p
9	2848.43	35096.8	6.7	9	5s ² 5p5d	³ F ₂ -5s ² 5p4f
11	2875.15	34770.6	{ 1.2		5s ² 5p5d	¹ P ₂ -5s ² 5p4f
			{ 0.2	6	5s ² 5p6s	³ D ₃
10	2927.64	34147.3	7.6	6	5s ² 5p6s	¹ P ₁ -5s ² 5p6p
3	2936.86	34040.1	0.3	6	5s ² 5p5d	³ P ₀
12	2937.57	34031.8	1.8	6	5s ² 5p6s	³ P ₀ -5s ² 5p6p
10	2954.93	33831.9	2.0	6	5s ² 5p6s	³ P ₁ -5s ² 5p6p
13	3015.09	33156.9	7.0	9	5s ² 5p5d	³ F ₂ -5s ² 5p4f
16	3077.71	32485.5	2.4		5s ² 5p5d	³ F ₄ -5s ² 5p4f
9	3109.49	32150.3	0.5	9	5s ² 5p5d	³ D ₃ -5s ² 5p4f
10	3145.57	31781.6	1.1	6	5s ² 5p6s	¹ P ₁ -5s ² 5p6p
13	3149.11	31745.8	5.8		5s ² 5p5d	³ F ₄ -5s ² 5p4f
5	3227.69	30973.0	3.1		5s ² 5p5d	³ F ₃ -5s ² 5p4f
6	3230.04	30950.5	0.6		5s ² 5p5d	¹ D ₂ -5s ² 5p4f
16	3305.96	30242.5	9.9	5	5s ² 5p5d	³ F ₂ -5s ² 5p4f
2	3309.11	30210.9	0.8		5s ² 5p5d	³ F ₄ -5s ² 5p4f
9	3312.92	30176.2	6.3	6	5s ² 5p6s	¹ P ₁ -5s ² 5p6p
16	3330.84	30013.9	4.2	5	5s ² 5p5d	³ F ₃ -5s ² 5p4f
9	3443.32	29033.4	3.4		5s ² 5p5d	³ F ₃ -5s ² 5p4f
8	3444.70	29021.8	2.6		5s ² 5p5d	¹ F ₃

continued

Table 1—continued

Intensity	λ [Å] (vac.)	σ [cm ⁻¹]		Ref.	Classification	
		Obs	Calc			
3	3556.98	28105.7	5.9		5s ² 5p5d	¹ P ₁ -5s ² 5p6p
11	3792.31	26361.7	2.4		5s ² 5p5d	³ D ₂ -5s ² 5p4f
12	3803.26	26285.8	5.7	5	5s ² 5p5d	³ P ₀ -5s ² 5p4f
14	3827.98	26116.1	6.3		5s ² 5p5d	³ F ₃ -5s ² 5p6p
6	3828.89	26109.8	10.0	6	5s ² 5p6s	³ P ₂ -5s ² 5p6p
3	3938.04	25386.2	6.7	6	5s ² 5p6s	³ P ₂ -5s ² 5p6p
6	3967.25	25199.3	9.5		5s ² 5p5d	³ D ₃ -5s ² 5p4f
8	3973.01	25162.7	3.1	5	5s ² 5p5d	³ P ₁ -5s ² 5p4f
4	4038.46	24754.9	4.5		5s ² 5p5d	³ D ₂ -5s ² 5p4f
10	4237.66	23591.3	1.6		5s ² 5p5d	³ D ₃ -5s ² 5p4f
15	4305.69	23218.5	9.1	5	5s ² 5p5d	³ P ₁ -5s ² 5p4f
12	4422.08	22607.5	7.6		5s ² 5p5d	¹ F ₃ -5s ² 5p4f
1	4515.67	22138.9	8.7	6	5s ² 5p6s	¹ P ₁ -5s ² 5p6p
15	4558.49	21931.0	0.9	12	5s5p ³	¹ P ₁ -5s ² 5p4f
12	4634.49	21571.3	1.7		5s ² 5p5d	¹ D ₂ -5s ² 5p4f
7	4769.94	20958.8	9.1	6	5s ² 5p6s	¹ P ₁ -5s ² 5p6p
9	4849.30	20615.8	5.9		5s ² 5p5d	³ P ₂ -5s ² 5p4f
15	4954.13	20179.6	9.7	5,9	5s ² 5p5d	¹ F ₃ -5s ² 5p4f
12	5007.80	19963.3	3.8	5	5s ² 5p5d	¹ D ₂ -5s ² 5p4f
1	5151.90	19405.0	5.1		5s ² 5p6s	³ P ₂ -5s ² 5p4f
15	5159.08	19377.9	8.0	5	5s ² 5p5d	³ D ₃ -5s ² 5p4f
15	5260.19	19005.4	5.9	5	5s ² 5p5d	³ D ₂ -5s ² 5p4f
16	5352.92	18676.2	6.2	5	5s ² 5p5d	³ P ₂ -5s ² 5p4f
15	5394.62	18531.9	1.8	5	5s ² 5p5d	³ D ₁ -5s ² 5p4f
9	5602.83	17843.2	3.0		5s ² 5p5d	³ D ₃ -5s ² 5p4f
2	5899.11	16947.0	7.3		5s ² 5p5d	³ F ₄ -5s ² 5p4f
12	5955.67	16786.1	6.6	5	5s ² 5p5d	¹ P ₁ -5s ² 5p4f
1	6653.85	15024.7	4.7		5s ² 5p6s	¹ P ₁ -5s ² 5p6p

Notes: Int. Observed line intensity from visual estimation.

λ (Å): Observed wavelength in Å.

vac. Values of the observed wavelength in vacuum.

air Values of the observed wavelength in air.

σ (cm⁻¹): Wave number in cm⁻¹.

Obs. Observed wave number.

Calc. Calculated wave number from the level values given in Table 2 and the work of Tauheed *et al.*⁸ by the Ritz combination principle.

Ref: References of previously published lines.

Classification: Level designations for the transition including configuration, term and total angular momentum.

levels involved. We present in total, 14 classified laser lines, 11 more than were observed in the work of Larsson *et al*⁹ using collision-based spectroscopy.

We present in Table 2 the adjusted energy level values corresponding to the analysed configurations. We have revised the energy values corresponding to the 5s²5p4f ³D₁ and ³G₅ levels reported in the work of Larsson *et al*,⁹ and changed them to 209310.0 and 202281.5 cm⁻¹, respectively. We also present in Table 2 two new energy levels values in 207366.7 and 216745.6 cm⁻¹ that correspond to the levels 5s²5p4f ³D₂ and ¹D₂, respectively, in order to complete all the energy level of this configuration. Our fitted values for these levels are also in agreement with the predicted values showed in the work of Larsson *et al*.⁹ In the same table is presented also the new value in 169799.4 cm⁻¹ for the 5s²5p5d ³F₄ energy level reported in the same reference. This new value is supported by four new classified lines and by the line in 3077.71 Å corresponding to the 5s²5p5d ³F₄-5s²5p4f ³G₅ transition.

Table 2. Energy level values of Xe V

Designation	E_{Obs} (cm $^{-1}$)	E_{Calc} (cm $^{-1}$)	Percentage composition
$5s^25p^2$	$(^3P)^3P_0$	0.0	15.3
	$(^3P)^3P_1$	9291.8	9300.8
	$(^3P)^3P_2$	14126.7	14150.2
	$(^1D)^1D_2$	28411.2	28387.8
	$(^1S)^1S_0$	44470.4	44476.1
	$(^2P)^3D_1$	228064.9	228484.3
	$(^2P)^3D_2$	235178.9	234968.0
	$(^2P)^3D_3$	246208.0	246101.9
	$(^2P)^3P_0$	233999.3	233755.0
	$(^2P)^3P_1$	234459.6	234452.7
$5s^25p6p$	$(^2P)^3P_2$	244821.3	244805.1
	$(^2P)^3S_1$	247810.4	247814.9
	$(^2P)^1D_2$	250557.2	251227.1
	$(^2P)^1P_1$	243216.5	242881.6
	$(^2P)^1S_0$	259642.3	259465.7
	$(^2P)^3G_3$	186746.7	186752.3
	$(^2P)^3F_4$	190644.7	190712.6
	$(^2P)^3G_5$	202281.8	202517.7
	$(^2P)^3F_2$	191603.5	191268.2
	$(^2P)^3F_3$	189663.8	189876.0
$5s^25p4f$	$(^2P)^3G_4$	201545.2	201559.1
	$(^2P)^3D_1$	209310.7	209139.1
	$(^2P)^3D_2$	207366.7	207204.4
	$(^2P)^3D_3$	205758.8	205913.2
	$(^2P)^1G_4$	214317.7	214683.1
	$(^2P)^1F_3$	200010.2	199883.4
	$(^2P)^1D_2$	216745.6	216483.3
	$(^4S)^5S_2$	92182.8	92202.7
	$(^2D)^3D_1$	115286.3	115276.0
	$(^2D)^3D_2$	116097.0	116021.0
$5s5p^3$	$(^2D)^3D_3$	119919.0	119931.7
	$(^2P)^3P_0$	133408.1	133406.5
	$(^2P)^3P_1$	134575.2	134536.5
	$(^2P)^3P_2$	134702.7	134737.0
	$(^4S)^3S_1$	155518.1	155545.3
	$(^2P)^1P_1$	169672.6	169435.1
	$(^2P)^1D_2$	145807.0	145976.2
	$(^2P)^3F_2$	156506.8	156462.7
	$(^2P)^3F_3$	160630.4	160663.5
	$(^2P)^3F_4$	169799.4	169759.2
$5s^25p5d$	$(^2P)^3D_1$	173071.7	173220.1
	$(^2P)^3D_2$	181004.3	181050.6
	$(^2P)^3D_3$	182167.2	181959.0
	$(^2P)^3P_0$	183025.2	183059.3
	$(^2P)^3P_1$	184147.6	184139.3

continued

Table 2—continued

Designation	E_{obs} (cm $^{-1}$)	E_{Calc} (cm $^{-1}$)	Percentage composition
(2P) 3P_2	170987.6	171018.9	48 + 25 (2P) 3D + 10 (2P) 1D + 12 5s5p 3 (2D) 1D
(2P) 1F_3	194138.0	194190.1	90 + 8 (2P) 3D
(2P) 1D_2	185795.0	185803.9	15 + 43 (2P) 3P + 22 (2P) 3D + 10 5s5p 3 (2D) 1D + 7 5s5p 3 (2P) 3P
(2P) 1P_1	199959.0	200018.4	55 + 13 p 3 (2P) 1P + 24 5s 2 5p6s (2P) 3P
5s 2 5p6s	(2P) 3P_0	194033.1	99
	(2P) 3P_1	194232.9	57 + 26 (2P) 1P + 10 5s 2 5p5d (2P) 1P
	(2P) 3P_2	209068.9	100
	(2P) 1P_1	213040.2	73 + 18 (2P) 3P + 8 5s 2 5p5d (2P) 1P

Notes:

Designation: Level designation, including configuration, parentage, term and total angular momentum.

E_{obs} : Experimental energies in cm $^{-1}$. The energy values for the even parity are based on odd parity levels given in the work of Tauheed *et al.*⁸ and wave numbers of Table 1 weighted according to their estimated uncertainties.

E Calc: Values obtained from the least-squares fitting.

Percentage Composition: Percentage composition of the experimental energy level. Percentages lower than 5% are omitted.

In accordance with our wavelength values, we classified the laser line in 3305.96 Å as a 5s 2 5p5d 3F_2 –5s 2 5p4f 3G_3 transition, rejecting the previous classification reported by Larsson *et al.*⁹ We also report in this work a new value in 259642.3 cm $^{-1}$ for the 5s 2 5p6p 1S_0 energy level.

The level values were determined in an iterative procedure where the wavenumbers of the observed lines were weighted according to their estimated uncertainties.

Theoretical predictions of the energy levels have been used in our analysis of the spectra. Relativistic Hartree–Fock calculations were used to obtain the energy parameters with the computer code developed by Cowan.¹² Predictions of the energy levels are then made by diagonalizing the energy matrices. A least-squares fit of the parameters to the observed levels was carried out involving configuration interactions.

In Table 3 we present the parameters obtained in the LSF calculations for the 5s 2 5p 2 , 5s 2 5p6p and 5s 2 5p4f configurations with their corresponding Hartree–Fock values and scaling factors. We have included the 5p 4 configuration in this calculation because the 5s 2 5p 2 configuration structure is affected through the s 2 –p 2 interaction and the standard deviation was reduced from 368 to 328 cm $^{-1}$. The inclusion of the 5s 2 5p6p 1S_0 level made possible to let free the G 2 (5p6p) parameter, which was fixed in the table of energy parameters reported by Gallardo *et al.*⁶

All the configuration interaction integrals were fixed at their Hartree–Fock values in the calculation except the 5s 2 5p 2 –5p 4 integral that it was fixed at 85% of its Hartree–Fock value. Hereby we achieved better results in the least-squares fit and the standard deviation was slightly reduced. These configurations structures are also affected by the interaction with the experimental unknown 5s5p 2 5d configuration, but its inclusion in this calculation significantly increases the standard deviation.

The fitted values for the odd configurations were calculated including configuration Rydberg series interaction and the new experimental value for the 5s5p 2 5d 3F_4 energy level. The results were in general accordance with the previously published values,⁸ except that our calculation shows a small reduction of the standard deviation value, considering that we were using different scaled values for the interaction integrals.

Table 3. Energy parameters for the $5s^25p^2$, $5s^25p6p$, $5s^25p4f$ and $5p^4$ configurations of Xe V

Configuration	Parameter	HFR value (cm ⁻¹)	Fitted value (cm ⁻¹)	Fit/HF
$5s^25p^2$	E_{av}		21 403	
	$F^2(5p,5p)$	55 554	$50\ 278 \pm 1007$ – 143 (FIX) ^a	0.905 ± 0.018
	$\alpha(5p,5p)$			
	ζ_{sp}	9095	9593 ± 202	1.055 ± 0.022
$5s^25p6p$	E_{av}	225 119	$242\ 484 \pm 113$	1.077 ± 0.001
	ζ_{sp}	9793	8739 ± 167	0.892 ± 0.017
	ζ_{ep}	2586	2594 ± 213	1.003 ± 0.082
	$F^2(5p,6p)$	19 780	$17\ 839 \pm 902$	0.902 ± 0.046
	$G^0(5p,6p)$	4267	3503 ± 149	0.821 ± 0.035
	$G^2(5p,6p)$	5648	5846 ± 809	1.035 ± 0.143
$5s^25p4f$	E_{av}	187 936	$200\ 751 \pm 96$	1.068 ± 0.001
	ζ_{sp}	8858	9330 ± 191	1.053 ± 0.021
	ζ_{4f}	168	168 (FIX)	1.000
	$F^2(5p,4f)$	48 022	$39\ 199 \pm 1173$	0.816 ± 0.024
	$G^2(5p,4f)$	37 402	$29\ 201 \pm 825$	0.781 ± 0.022
	$G^4(5p,4f)$	26 627	$18\ 663 \pm 1123$	0.700 ± 0.042
$5p^4$	E_{av}	249 988	249 988 (FIX)	1.000
	$F^2(5p,5p)$	55 559	55 559 (FIX)	1.000
	$\alpha(5p,5p)$		0 (FIX)	
	ζ_{sp}	8994	8994 (FIX)	1.000
<i>Configuration interaction integrals</i>				
$5s^25p^2-5s^25p6p$	$R^0(5p5p, 5p6p)$	1851	1851 (FIX)	1.000
$5s^25p^2-5s^25p6p$	$R^2(5p5p, 5p6p)$	8658	8658 (FIX)	1.000
$5s^25p^2-5s^25p4f$	$R^2(5p5p, 5p4f)$	– 43 668	– 43 668 (FIX)	1.000
$5s^25p^2-5p^4$	$R^4(5s5s, 5p5p)$	72 662	55 680 (FIX)	0.766
$5s^25p6p-5s^25p4f$	$R^2(5p6p, 5p4f)$	– 5314	– 5314 (FIX)	1.000
$5s^25p6p-5s^25p4f$	$R^2(5p6p, 4f5p)$	– 6025	– 6025 (FIX)	1.000
σ^b			328	

Notes:

Configuration: Configurations involved in the calculation.

Parameter: Energy parameters.

HFR: Hartree–Fock values in cm⁻¹ of the energy parameters.Fitted value: Fitted value of the energy parameters: a, The α value is taken from the work of Tauheed *et al.*⁸.

Fit/HF: Rate between fitted and Hartree–Fock energy parameter values.

Configuration interaction integrals: Values of the Configuration interaction integrals for the involved configurations.

b, The rms. deviation of the fit.

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REFERENCES

1. Arcimowicz, B., Yoshi, Y. N. and Kaufman, V., *Can. J. Phys.*, 1989, **67**, 572.
2. Tauheed, A., Yoshi, Y. N. and Kaufman, V., *J. Phys. B*, 1991, **24**, 3701.
3. Tauheed, A., Yoshi, Y. N. and Kaufman, V., *Phys. Scr.*, 1991, **44**, 579.
4. Tauheed, A. and Yoshi, Y. N., *Phys. Scr.*, 1992, **46**, 403.
5. Duchowicz, R., Schinca, D. and Gallardo, M., *IEEE J. Quantum Electron.*, 1994, **30**, 155.
6. Gallardo, M., Raineri, M. and Reyna Almandos, J.G., *Z. Phys. D*, 1994, **30**, 261
7. Pinnington, E. H., Gosselin, R.N., Ji, Q., Kernahan, J. A. and Guo, B., *Phys. Scr.*, 1992, **46**, 40.
8. Tauheed, A., Yoshi, Y. N., Kaufman, V., Sugar, J. and Pinnington, E. H., *J. Opt. Soc. Am. B*, 1993, **10**, 561.
9. Larsson, M. O., Gonzales, A. M., Hallin, R., Heijkenskjöld, F., Nyström, B., O'Sullivan, G., Weber, C. and Wännström, A., *Phys. Scr.*, 1996, **53**, 317.
10. Valero, F. P. J., *J. Opt. Soc. Am.*, 1968, **58**, 1048.
11. Hoffmann, V. and Toschek, P., *IEEE J. of Quantum Electron.*, 1970, **6**, 757.
12. Cowan, R.D., *The Theory of Atomic Structure and Spectra*. University of California Press, Berkeley, 1981.